Space debris

Laboratory of Spacecraft Environment Interaction Engineering (La SEINE) Kyushu Institute of Technology, Japan

Introduction to space debris

The source of space debris could be expired satellites, spent rocket upper stages, fragments from explosions and collisions, paint flakes, chunks of slag from solid rocket motors, remnants of old science experiments and a variety of small particles.

Three categories of space debris, depending on their size: Category I (<1cm) make significant damage to vulnerable parts of a satellite, shielding

Category II (1-10 cm) seriously damage or destroy a satellite in a collision, no effective shielding

Category III (>10cm) destroy a satellite in a collision, can be tracked(in GEO >1m), evasive maneuver

Classification of debris



The first artificial space debris is the first satellite launched into space was the Soviet Sputnik-1. This was placed into low Earth orbit on 4 October 1957.



Exhaust

Products



& Mission Operations

Human Missions

http://www.spaceacademy.net.au/watch/debris/gsd/gsd.htm#class

Radioactive space debris:

-Fragments of nuclear propulsion systems;

-Individual SC with nuclear propulsion systems which operational cycle has been finished

-Operating SC equipped with one or another nuclear installation

Methods are proposed to detect and identify space debris radioactive objects



Amount of orbital debris

- currently more than 20,000 objects are tracked and kept in a catalog

- space debris number is much more than catalog

| Debris Size | 0.1-1cm | 1-10cm | >10cm | | |
|---|-------------|---------|--------|--|--|
| Total Number at all altitudes | 150 million | 780,000 | 23,000 | | |
| Debris in Low-Earth Orbit | 20 million | 400,000 | 15,000 | | |
| Estimated amount of orbital debris, by size | | | | | |

Distribution of LEO debris



Evolution of total trackable Low-Earth Obit (LEO) object



Marshall H. Kaplan, Survey of Space Debris Reduction Methods, AIAA SPACE 2009 Conference & Exposition 14 - 17 September 2009, Pasadena, California, AIAA 2009-6619 The number of space debris increases very

Collision of Iridium33 and Cosmos2251

In an unprecedented space collision, a commercial communications satellite (IRIDIUM 33) and a defunct **Russian satellite (COSMOS** 2251) impacted each other on February 9th, 2009 above Northern Siberia, creating a cloud of debris. Till now, over 1719 large fragments have been observed.

VIDEO link https://www.youtube.com/watch?v=dtT3pTh_q -8



Spread of collision debris orbital planes



Debris observation

Space object catalogues are limited to a larger objects, typically greater than 10 cm in low Earth orbits and greater than 1m at geostationary altitude

Ground based telescopes and radars can detect: -LEO debris down to a few mm in size -GEO debris down to 10 cm in size

In-situ impact detector (on-board spacecraft) can sense object down to a few micrometers in size While telescopes are mainly suited for GEO and high-altitude debris observations, radars are advantageous in the low-Earth orbit (LEO) regime, below 2000 km.

Radar facilities for debris observation

| Country | Organization | Facility | Type | Primary operation mode | Configuration | Field of view | Wave- length (m) | Sensitivity (diameter) (m) | Status |
|-------------------------------|---------------------|------------|-----------------|------------------------------|-----------------------|---------------------|------------------------|----------------------------------|------------------------|
| Germany | FGAN | TIRA | Dish | Mixed | Monostatic | 0.5 | 0.23 | 0.02 at 1,000 km | Operational |
| Germany | MPIfR | Effelsberg | Dish | Stare | Bistatic with TIRA | 0.16 | 0.23 | 0.009 at 1,000 km | Experimental |
| Japan | Kyoto University | MU radar | Phased array | Stare | Monostatic | 3.7 | 6.4 | 0.02 at 500 km | Operational |
| Japan | ISAS | Uchinoura | Dish | Mixed | Bistatic | 0.4 | 0.13 | 0.02 at 500 km | Experimental |
| Japan | ISAS | Usuda | Dish | Mixed | Bistatic | 0.13 | 0.13 | 0.02 at 500 km | Experimental |
| Ukraine/Russian Federation | | Evpatoria | Dish | Stare | Bistatic | 0.1 | 0.056 | 0.003 at 1,000 km | Developmental |
| United States | NASA/ NSF | Arecibo | Dish | Stare | Bistatic | 0 | 0.13 | 0.004 at 575 km | One-time experiment |
| United States | NASA/ DoD | Haystack | Dish | Stare | Monostatic | 0.1 | 0.03 | 0.006 at 1,000 km | Operational |
| United States | NASA/ DoD | HAX | Dish | Stare | Monostatic | 0.1 | 0.02 | 0.05 at 1,000 km | Operational |
| United States | NASA | Goldstone | Dish | Stare | Bistatic | 0 | 0.035 | 0.002 at 500 km | Operational |
| United States | DoD | TRADEX | Dish | Mixed | Monostatic | 0.61/ 0.30 | 0.23/ 0.10 | 0.03 at 500 km | Operational |
| | | | | | | | | | |

Table 1. Radar facilities for debris observation

The probability of collision in the radial direction

The uncertainly the radial position is defined by a onedimensional normal distribution with the form:

$$\rho(r) = \frac{1}{2\sqrt{\pi}\sigma_r} \exp\left\{\frac{-(r-\Delta h)^2}{2\sigma_r^2}\right\}$$

Where Δh is the difference in altitude of the two objects, and the relative error in position in the radial direction is given by:

$$\sigma_r^2 = \sqrt{\sigma_{1r}^2 + \sigma_{2r}^2}$$



Then the density function describing the probability of collision can be given by

$$F(r) = \frac{1}{2\sqrt{\pi}\sigma_r} \int_{-\infty}^r \exp\left\{\frac{-(r-\Delta h)^2}{2\sigma_r^2}\right\} \mathrm{d}r$$

The probability of collision in the radial direction

It can be substituted by an approximate expression with the form:

$$F(r) = \frac{\exp\left[a(r - \Delta h)\right]}{1 + \exp\left[a(r - \Delta h)\right]}, \quad \text{where } a = \frac{4}{\sqrt{2\pi}\sigma_r}$$

The result from this approximation is sufficiently accurate and does not introduce error greater than 2%.

Denoting the combined size of the objects in the radial direction as r_a, the probability of collision caused by difference in altitude can be expressed as:

$$P_R = F(x) \mid_{-\infty}^{r_a} - F(x) \mid_{-\infty}^{-r_a} = \frac{\exp\left[\frac{4(r_a - \Delta h)}{\sqrt{2\pi\sigma_r}}\right]}{1 + \exp\left[\frac{4(r_a - \Delta h)}{\sqrt{2\pi\sigma_r}}\right]} - \frac{\exp\left[\frac{4(-r_a - \Delta h)}{\sqrt{2\pi\sigma_r}}\right]}{1 + \exp\left[\frac{4(-r_a - \Delta h)}{\sqrt{2\pi\sigma_r}}\right]}$$

Mean time between impacts

Table 5. Mean time between impacts on a satellite with a cross-section area of 10 square metres

| Height of circular orbit | Objects 0.1-1.0 cm | Objects 1-10 cm | Objects >10 cm |
|--------------------------|--------------------|-------------------|----------------|
| 500 km | 10-100 years | 3,500-7,000 years | 150,000 years |
| 1,000 km | 3-30 years | 700-1,400 years | 20,000 years |
| 1,500 km | 7-70 years | 1,000-2,000 years | 30,000 years |

Orbital Debris Distribution

- Largest portion (2/3) of orbital debris is concentrated in LEO
- Only 6% of Earth orbiting objects are operational payloads



Altitude Distribution in LEO



http://www.surrey.ac.uk/ssc/research/space_vehicle_control/deorbitsail/files/gsf_keynote_2012.pdf

Orbital Debris projection



Catalogued Objects in Orbit

Catalogued Objects in Orbit as of October 2012



In almost 50 years of space activities, more than 4900 launches have placed some 6600 satellites into orbit, of which about 3600 remain in space; only a small fraction - about 1000 - are still operational today.

Active debris removal method

Removing debris from LEO can make the LEO environment safe for the future space activities. And clearing GEO will keep the GEO orbit resources available.

| | Catalog I (Size < 1cm) | Catalog II (Size 1cm~10cm) | Catalog III (Size >10cm) |
|-----------------------------------|---|---------------------------------------|--|
| LEO orbit 160km- 2,000km | Space-based Magnetic Field Generator Sweeping/Retarding Surface Space-based Laser | Ground-/Air- /Space-based Laser | Drag Augmentation Device Magnetic Sail Momentum Tethers Electrodynamic Tethers Capture/Orbital Transfer Vehicle (Space Shutter) |
| GEO orbit About 35,000km | | | Solar Sail Momentum Tethers Capture/Orbital Transfer Vehicle (Using Net or Tentacles) |

VIDEO LINK- http://www.star-tech-inc.com/id121.html

Space debris removal method using electrostatic force in space plasma

Debris removal method for small diameter less than 1 mm. The debris must be decelerated in order to descend their orbit. This method employs the deceleration of debris by means of electrostatic force. Debris must be charged electrically with respect to space (usually have same potential with LEO plasma)



Space debris removal method using electrostatic force in space plasma

The space potential increases from plasma potential to bias potential in negative sheath area with approaching the net. The capacitance between debris and space are

$$C_d = 4\pi\varepsilon_0 a$$

The debris has a negative charge Qd

$$Q_d = C_d \left(V_s - V_f \right) = 4 \pi \varepsilon_0 a \left(V_s - V_f \right)$$

After passing through the net, the negative charged debris is pulled back in the direction of the bias net due to electrostatic force in the negative sheath. The debris loses the energy U_{dc} .

$$U_{dc} = Q_d V_{bias} = 4\pi\varepsilon_0 a V_{bias}^2$$

Space debris removal method using electrostatic force in space plasma

During passing the biased net, it is assumed that the debris loses the energy U_{dc} , the debris velocity decreases from v_0 to v_1 , and the altitude descends from h_0 to h_2

$$\frac{1}{2}mv_{1}^{2} - G\frac{Mm}{R+h_{0}} = \frac{1}{2}mv_{2}^{2} - G\frac{Mm}{R+h_{2}}$$
$$\frac{mv_{2}^{2}}{R+h_{2}} = G\frac{Mm}{(R+h_{2})^{2}}$$
$$\frac{1}{2}mv_{0}^{2} = \frac{1}{2}mv_{1}^{2} + U_{dc}$$
$$h_{2} = \frac{GM}{\frac{GM}{R+h_{0}} + \frac{6\varepsilon_{0}V_{bias}}{\rho a^{2}}} - R$$

Atmospheric drag

$$\mathbf{a}_{D} = \frac{\mathbf{D}}{m} = -\frac{1}{2} \frac{C_{D}A}{m} \rho \left| \mathbf{v}_{rel} \right| \mathbf{v}_{rel}$$

- Speed relative to the atmosphere \mathbf{v}_{rel}
- **D** C_D Drag Coefficient:

▶ Rotation, winds

depends on attitude

 $> \sim 15\%$ error

- difficult to measure
- $\triangleright C_D \sim 2 2.4$ (1-4)

- A Frontal area
- ρ Atmospheric density:

■ $\beta = \frac{m}{C_D A}$ Ballistic coefficient: (β ↑, $\mathbf{a}_D \downarrow$)

Some authors use the opposite form: $BC = \frac{C_D A}{C_D A}$

Atmospheric drag



http://ocw.upm.es/ingenieria-aeroespacial/modeling-the-space-environment/contenidos/material-de-clase/mse06_atmos.pdf

Atmospheric drag

SATELLITE ORBITAL DECAY

Calculating decay rates and orbital lifetimes of satellites in essentially circular orbits below 500 km altitude.

| Mass | 100.0 | kg | |
|--------|------------------------------|-----|----------------------------|
| Area | 1.0 | m^2 | |
| Initia | l Altitutde 500 | | km (range: 180km - 500 km) |
| Solar | Radio Flux (F10.7) 70 | | |
| Geom | agnetic Index 0 | | (Ap) |
| cal | c | | |

| TIME(days) | HEIGHT(km) | PERIOD(mins) | MEAN MOTION(rev/day) | DECAY(rev/day^2) |
|------------|------------|--------------|----------------------|------------------|
| 0 | 500 | 94.58 | 15.22 | 0 |
| 1364.09 | 489.99 | 94.39 | 15.25 | 0 |
| 2608.39 | 479.99 | 94.19 | 15.28 | 0 |
| 3605.79 | 469.99 | 93.98 | 15.32 | 0 |
| 4403.09 | 459.99 | 93.78 | 15.35 | 0 |
| 5038.8 | 449.99 | 93.57 | 15.38 | 0 |
| 5544.3 | 439.99 | 93.36 | 15.42 | 0 |
| 5945.2 | 429.99 | 93.16 | 15.45 | 0 |
| 6262.2 | 419.99 | 92.95 | 15.49 | 0.0001 |
| 6512.3 | 409.99 | 92.75 | 15.52 | 0.0001 |
| 6709.1 | 399.99 | 92.54 | 15.55 | 0.0001 |
| 6863.5 | 389.99 | 92.34 | 15.59 | 0.0002 |
| 6984.3 | 379.99 | 92.13 | 15.62 | 0.0003 |
| 7078.6 | 369.99 | 91.93 | 15.66 | 0.0004 |
| 7152 | 359.99 | 91.73 | 15.69 | 0.0005 |
| 7209 | 349.98 | 91.52 | 15.73 | 0.0007 |

http://www.lizard-tail.com/isana/lab/orbital_decay/

Solar sail





| Mass (kg) | 3 | | 150 | | 500 | |
|-----------------------|---|---|--|--|---|--|
| Area (m²) | 0.03 | 25 | 0.8 | 25 | 1.5 | 25 |
| Initial altitude (km) | t _D (<u>years</u>) without sail | t _D (<u>days</u>) with sail | t _D (<u>years)</u> without sail | t _D (<u>years</u>) with sail | t _D (<u>years</u>) without sail | t _D (<u>years</u>) with sail |
| 600 | 26.5 | 39 | 42.4 | 2.3 | 73.3 | 5.3 |
| 650 | 48.1 | 75 | 85.0 | 3.7 | 146.5 | 13.6 |
| 700 | 88.5 | 132 | 159.9 | 6.7 | 282.7 | 24.3 |
| 750 | 160.5 | 190 | 299.7 | 14.1 | 521.5 | 37.9 |

Satellite deorbit system based on electric



propulsion Weight: 6 gr (thruster head) 32 gr (all system) Size : ϕ 13 mm x 21 mm Applied voltage = 300 V...800 V (thruster head) $I_{sp} = 1200[s]$ ANODE Impulse bit ~ 1.4 μ Ns INSULATOR CATHODE 99 21 ANODE CATHODE INSULATOR SATEL ITE BODY



Principle of Vacuum Arc Thruster



Characteristics:

- Solid propellant;
- Metal vapor is jetted;
- 300 V direct drive;
- Firing by the micro discharge in plasma environment.

Principle of Vacuum Arc Thruster



Characteristics:

- Solid propellant;
- Metal vapor is jetted;
- 300 V direct drive;
- Firing by the micro discharge in plasma environment.

Ion concentration



RAM side compression region would be:

$$\begin{aligned} v_{th} &= \sqrt{\frac{2 \cdot k \cdot T}{m_p}} \\ k &= 1.38 \cdot 10^{-23} \\ T &= 2327(K) \\ m_p(O) &= \frac{16 \cdot 10^{-3}}{6.23 \cdot 10^{23}} = 2.6 \cdot 10^{-26} (kg) \\ v_{th} &= 1572(\frac{m}{\text{sec}}) \\ n_{O^+} &= 1.3 \cdot 10^{11} (m^{-3}) \\ n_0 &+ n_r = n_0 \cdot (1 + \frac{V_0}{v_{th}}) = 1.3 \cdot 10^{11} \cdot (1 + \frac{7800}{1572}) = 7.75 \cdot 10^{11} (m^{-3}) \end{aligned}$$

Behind the satellite body:

$$n \approx n_0 \exp\left[-\left(\frac{V_0}{v_{th}}\right)^2 \cdot \left(\frac{R_0}{r}\right)^2\right]$$
$$R_0 = 0.05(m)$$
$$r = 0.05(m)$$

$$n \cong 1.3 \cdot 10^{11} \exp\left[-\left(\frac{7800}{1572}\right)^2 \cdot \left(\frac{0.05}{0.05}\right)^2\right] = 0(m^{-3})$$

Satellite deorbit system based on electric propulsion



$$\frac{mv_1^2}{2} - G\frac{Mm}{R+h} = \frac{mv_2^2}{2} - G\frac{Mm}{R+H}$$

$$F = \frac{1}{2} \cdot C_d \cdot A \cdot \rho \cdot v_1^2 - atmospheric_drag$$

$$v_1 = \sqrt{G \frac{M}{R+h}}$$

$$mv_1 - mv_2 = I_{bit}$$

$$v_2 = v_1 - \frac{I_{bit} \cdot N_p}{m};$$

$$m - satellite _mass$$

$$M = 6 \cdot 10^{24} [kg]$$

$$R = 6.37 \cdot 10^{6} [m]$$

$$G = 6.67 \cdot 10^{-11} [m^{3}s^{-2}kg^{-1}]$$

$$h = 600 [km]$$

$$H = 400 [km]$$

$$N_{p} = \frac{m}{I_{bit}} \left[\sqrt{G \frac{M}{R+h}} - \sqrt{2 \cdot G \frac{M}{R+h}} - G \frac{M}{R+H} \right]$$

Satellite deorbit system based on electric propulsion N_p





Satellite deorbit system based on electric propulsion





| deorbit time, years | without magnet | | NS middle | NS section |
|---------------------------|-------------------|------|--------------|---------------|
| 300 | | 863 | 705 | 783 |
| 400 | | 651 | 525 | 539 |
| 500 | | 123 | 94.9 | 94.5 |
| 600 | | 50.4 | 38.5 | 36.1 |
| 700 | | 43.9 | 32.3 | 30.8 |
| 800 | | 51.3 | 36.3 | 36.1 |





| deorbit time, | without | NS | |
|------------------|---------|--------|------------|
| years | magnet | middle | NS section |
| 300 | 137 | 93.4 | 79.1 |
| 400 | 54.9 | 38.6 | 36.4 |
| 500 | 28.3 | 20.4 | 20.8 |
| 600 | 14.3 | 10.5 | 11.4 |
| 700 | 3.63 | 2.69 | 3.07 |
| 800 | 2.41 | 1.8 | 2.15 |



Debris

Evolution

- Euroconsult forecast for next 10 years shows 400 out of 1200 anticipated launches will be in LEO – only includes satellites
 > 50kg
 To have a sustainable LEO population requires:
- Implementation of commonly adopted mitigation measures
- Active Debris Removal (ADR) of 5 objects or more...

25 Year Deorbiting Requirement

Orbital debris mitigation guidelines have been debated in great detail in the space community.
 In February 2007 and after a multi-year effort the United Nations' Committee on the Peaceful Uses of Outer Space (COPUOS) adopted a set of space debris mitigation guidelines which includes a 25 year deorbit requirement LEO.

• The guidelines were accepted by the COPUOS in June 2007 and endorsed by the United Nations in January 2008.

• To become the law in many countries in Europe

Thank you for your attention